RESIDENT ARCHIVE SERVICES OF THE YOHKOH LEGACY DATA ARCHIVE

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ABSTRACT

Yohkoh Legacy data Archive (YLA) is one of the first group of the Resident Archives (RAs) selected for funding for NASA's Virtual Observatories for Heliophysics Data program. YLA provides the best corrected data set of solar X-ray images and spectra from the Yohkoh satellite with user-friendly web interface. As a RA, we take responsibility to keep our products well maintained and provide access to the archive. In addition, we launched the `E-consultant service`, e-mail based support to individual users regarding the data handling, as our attempt to bolster the access and use from a wide range of communities.

Keywords: Resident Archive, Virtual Observatories, Yohkoh, Yohkoh Legacy data Archive, YLA, Solar corona, X-ray images

1 THE YOHKOH MISSION

Yohkoh is a Japan/US/UK mission investigating the high-energy phenomena in the solar corona. Yohkoh was in operation for more than ten years (September 1991 through December 2001), which covers almost the whole solar activity cycle. Yohkoh carries four scientific instruments: Soft X-ray Telescope (SXT), Hard X-ray Telescope (HXT), Bragg Crystal Spectrometer (BCS), and Wide Band Spectrometer (WBS). Technical details of those instruments are found in the following papers; Tsuneta, et al, 1991, Kosugi, et al., 1991, Culhane, et al, 1991, and Yoshimori, et al, 1991. Yohkoh brought considerable progress in understanding the dynamic aspects of the solar corona:

- Discovering new active phenomena, e.g. transient brightenings (Shimizu, et al., 1994), X-ray jets (Shimojo, et al., 1996), and sigmoid eruptions (Canfield, et al., 1999).
- Confirming magnetic reconnection as the fundamental mechanism of energy release in flares, by revealing such evidences as cusp-shaped loops (Tsuneta, 1996), separate hard-X-ray and soft-X-ray sources (Masuda, et al., 1994), X-ray plasmoids (Ohyama & Shibata, 1998), supra arcade structures (McKenzie & Hudson, 2001) and velocity fields (Culhane, et al., 1994).

Yohkoh data are valuable as a steady source of the coronal images over the whole solar cycle. In particular, due to the fact that Yohkoh is the only coronal imager in orbit in early 1990s, the data for these period are unique and invaluable. The data processing environment developed for the Yohkoh project forms the foundation of the current data handling system, integrated as SolarSoft in the Interactive Data Language (IDL) and widely used in the solar physics community.

2 THE YOHKOH LEGACY DATA ARCHIVE

The Yohkoh Legacy data Archive project was kicked off in July 2002, with the following founding philosophy: Provide a fully-calibrated archive of the Yohkoh mission data, with easy access, so that any scientist can perform reliable analysis even when removed by distance and time from the core team of `Yohkoh experts`.

The main feature of YLA is in the improved quality of the soft X-ray images and high accumulation of value-added products and documentation:

Table 1. Basic data of the Yohkoh Legacy data Archive.

SXT products:

Level 0 weekly files in XDA format

Level 1 corrected data in XDA and FITS formats

Level 2 scientific composites (XDA, FITS), browse (movie) images (FITS), aspect sensor (white light)

corrected movie images (FITS), synoptic maps.

HXT, WBS, and BCS products:

Level 0 weekly files in XDA format

Level_1 N/A

Level 2 flare catalogs in HTML and ASCII formats

Documents: operation reports, calibration notes, analysis guides, the Science Nuggets, operational

procedures, SXT observation tables.

User interface: data search (YLA, VSO), SXT Movie Maker, SXT Quick Look, observing log inquiry.

Web URL: http://solar.physics.montana.edu/ylegacy

- 1. Providing corrected data set: We classify our products into 3 types, i.e., level 0, 1, and 2. The level_0 data are those of the most basic level, essentially reformatted raw telemetry data. The level_1 data are the best corrected version of the level_0 data. The further processed products, e.g., composite images, synoptic maps, event catalogs, are classified as level_2. Table 1 is the summary of contents of the YLA as of December 2008. During more than ten years of operation, Yohkoh had trouble in the SXT entrance filter and satellite attitude control system. The former caused significant increase of stray light level while the latter caused alignment and roll problem in SXT images. We have spent much time to correct those effects and improve the quality of our images, and a part of these efforts are still in progress.
- 2. Providing FITS files as well as the Yohkoh original format: Yohkoh data handling has long depended on the specific data format called XDA, which requires IDL software to handle. However, FITS format is now commonly used in many solar projects after Yohkoh. Providing FITS data therefore gives a common basis with recent data and improves manageability and accessibility of our products.
- 3. Development of the user-friendly web interface: We set up our web interface on a domain of Montana State University, Department of Physics. Figure 1 shows the top page of our archive. We developed our own data search, in which users can access by specifying the observed time range and optional parameters. YLA has linkage with the Virtual Solar Observatory (VSO) since 2005. Our level_0 and level_1 products are accessed through the VSO data search, while the accommodation of level_2 data to VSO is now in preparation. A quick-look service of our daily observation is available as SXT Quick Look (see Figure 2). We also provide a web interface for creating customized movies, called SXT Movie Maker.
- 4. Accumulation of value-added data and documents: Examples of our value-added products are our 2 kinds of flare catalogs. One is for the events detected with HXT/WBS instruments, and the other is for flares detected with BCS instrument (see Figure 3). The processing of the former catalog was initiated by the former HXT group in Japan (Sato, et al., 2006), while the latter is the contribution from the former BCS group in UK. Our documentation archive is rich in both amount and quality. It ranges from the records of daily operation to the short scientific articles.

The major part of the work has been achieved at Montana State University, with technical support from Lockheed Martin Solar and Astrophysics Laboratory. As already demonstrated in the above, we have data contributions from the former Yohkoh teams in Japan and the UK, and in some cases from individual researchers.



Figure 1. The front page of the Yohkoh Legacy data Archive web interface. The URL is http://solar.physics.montana.edu/ylegacy.

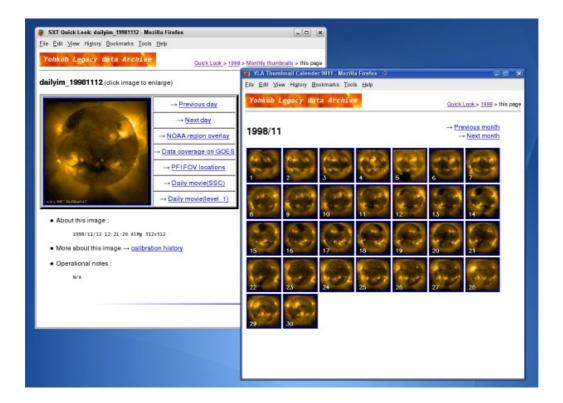


Figure 2. Sample pages from the SXT Quick Look service: a daily page (left) and a clickable monthly thumbnail images (right).

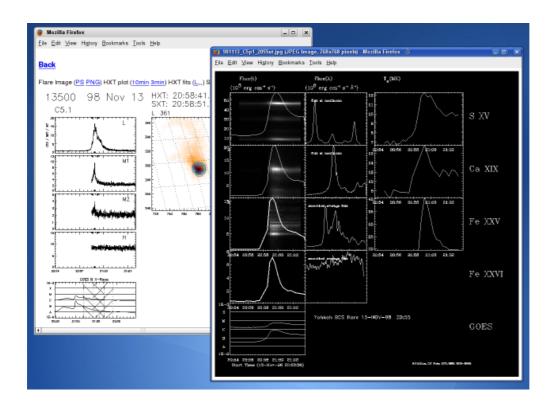


Figure 3. Sample pages of our flare catalogs: from the HXT/WBS flare catalog (left) and from the BCS flare catalog (right).

3 RESIDENT ARCHIVE SERVICES

Resident Archive is a new category of archives defined in NASA's Helio Physics Data Environment (HPDE) since 2007 (Fisher, 2007). Placing between Mission Archives and Final Archives, Resident Archives will provide continued access and maintenance to the post-mission data products optionally with expert support to the users. In 2008, YLA was selected for funding as one of the resident archives for HPDE's Virtual Observatories for Heliophysics (VxO) Data Program. Our proposed tasks as one of the resident archives are as follows.

- 1. Keep our products well maintained and avoid permanent loss.
- 2. Provide convenient and comfortable access to users.
- 3. Provide users individual assistance for accessing and handling data (E-consultant service).
- 4. Enhance the connection with the Virtual Solar Observatory (VSO) data search.
- 5. Improvement of analysis software, mainly to accommodate FITS users.
- 6. Preparing a 'transplant package' of our products for any group or individual who would like to have them.

Our greatest emphasis here is the E-consultant service. This service is provided in an effort to promote the use of our products. In analogy with the pharmacists at a drug store, we will assist users when they need support. We assume the consultation over the e-mail, via our contact address, *yla@solar.physics.montana.edu*. We will keep this service at least for two years since the fiscal year 2009 or as long as funded.

We are delighted to kick off our new services concurrently with the two anniversaries related to geo- and astrophysics; fifty years after the International Geophysical Year (1957-1958) and four hundred years after the first use of an astronomical telescope by Galileo Galillei in 1609 (celebrated as the International Year of Astronomy in 2009). Through the good maintenance and support services as a resident archive, we ambitiously aim to ensure the usefulness of the Yohkoh data over decades or even centuries, similar in spirit to the enduring accessibility and utility of the sunspot data recorded by Galillei in the 17th century. Although YLA is designed primarily to serve for the solar and terrestrial physics, many of our products have suitable quality for public use. Access from a wide range of communities is greatly encouraged.

4 REFERENCES

Canfield, R.C., Hudson, H.S., & McKenzie, D.E., (1999) Sigmoidal Morphology and Eruptive Solar Activity, *Geophysics Research Letters*, 26, 627-630.

Culhane, J.L., Hiei, E., Doschek, G.A., & et al., (1991) The Bragg Crystal Spectrometer for Solar-A, *Solar Physics*, 136, 89-104.

Culhane, J.L., Phillips, A.T., Inda-Koide, M., & et al., (1994) YOHKOH observations of the creation of high-temperature plasma in the flare of 16 December 1991, *Solar Physics*, 153, 307-336.

Fisher, R.R., (2007) NASA Heliophysics Science Data Management Policy, Retrieved December 19, 2008 from the World Wide Web: http://hpde.gsfc.nasa.gov/HPDP.html.

Freeland, S.L., (2000) SolarSoft, Retrieved December 19, 2008 from the World Wide Web: http://www.lmsal.com/solarsoft/sswdoc/index_menu.html.

Kosugi, T., Makishima, K., Murakami, T., & et al., (1991) The Hard X-ray Telescope for the Solar-A Mission, *Solar Physics*, 136, 17-36.

Masuda, S., Kosugi, T., Hara, H., Tsuneta, S., & Ogawara Y., (1994) A Loop-Top Hard X-ray Source in a Compact Solar Flare as Evidence for Magnetic Reconnection, *Nature*, 371/6497, 495-497.

McKenzie, D.E., & Hudson, H.S., (2001) Downflows and Structure above LDE Arcades: Possible Signatures of Reconnection?, *Proceedings of The University of Tokyo Symposium in 2000 on Magnetic Reconnection in Space and Laboratory Plasmas*, M. Hoshino, R. L. Stenzel, and K. Shibata, eds., Earth Planets Space, 53, 577-580.

Ohyama, M., & Shibata, K., (1998) X-Ray Plasma Ejection Associated with an Impulsive Flare on 1992 October 5: Physical Conditions of X-Ray Plasma Ejection, *Astrophysical Journal*, 499, 934-944.

Sato, J., Matsumoto, Y., Yoshimura, K., & et al. (2006) YOHKOH/WBS Recalibration and a Comprehensive Catalogue of Solar Flares Observed by YOHKOH SXT, HXT and WBS Instruments, *Solar Physics*, 236, 351-368.

Shimizu, T., Tsuneta, S., Acton, L.W., Lemen, J.R., Ogawara, Y., & Uchida, Y., (1994) Morphology of Active Region Transient Brightenings with the YOHKOH Soft X-ray Telescope, *Astrophysical Journal*, 422, 906-911.

Shimojo, M., Hashimoto, S., Shibata, K., Hirayama, T., Hudson, H.S., & Acton, L.W., (1996) Statistical Study of Solar X-Ray Jets Observed with the YOHKOH Soft X-Ray Telescope, *Publ. of the Astronomical Society of Japan*, 48, 123-136.

Tsuneta, S., Acton, L.W., Bruner, M.E., & et al., (1991) The Soft X-ray Telescope for the Solar-A Mission, *Solar Physics*, 136, 37-67.

Tsuneta S., (1996) Structure and Dynamics of Magnetic Reconnection in a Solar Flare, *Astrophysical Journal*, 456, 840-849.

Yoshimori, M., Okudaira, K., Hirasima, Y., & et al., (1991) The Wide Band Spectrometer on the Solar-A, *Solar Physics*, 136, 69-88.